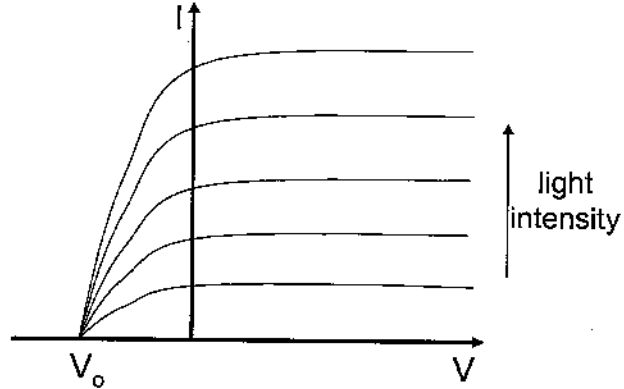
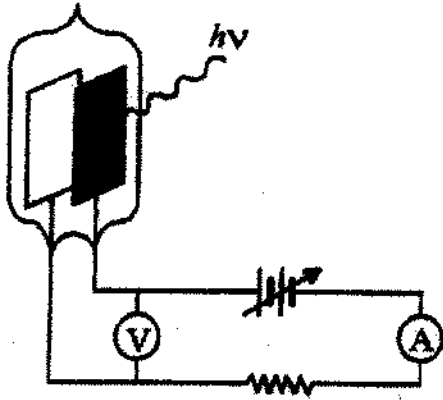


## Chapter 2 Solutions

### Prob. 2.1

(a&b) Sketch a vacuum tube device. Graph photocurrent  $I$  versus retarding voltage  $V$  for several light intensities.



Note that  $V_0$  remains same for all intensities.

(c) Find retarding potential.

$$\lambda = 2440 \text{ \AA} = 0.244 \mu\text{m} \quad \Phi = 4.09 \text{ eV}$$

$$V_0 = h\nu - \Phi = \frac{1.24 \text{ eV} \cdot \mu\text{m}}{\lambda(\mu\text{m})} - \Phi = \frac{1.24 \text{ eV} \cdot \mu\text{m}}{0.244 \mu\text{m}} - 4.09 \text{ eV} = 5.08 \text{ eV} - 4.09 \text{ eV} \approx 1 \text{ eV}$$

### Prob. 2.2

Show third Bohr postulate equates to integer number of DeBroglie waves fitting within circumference of a Bohr circular orbit.

$$r_n = \frac{4\pi\epsilon_0 n^2 \hbar^2}{mq^2} \quad \text{and} \quad \frac{q^2}{4\pi\epsilon_0 r^2} = \frac{mv^2}{r} \quad \text{and} \quad p_\theta = mvr$$

$$r_n = \frac{4\pi\epsilon_0 n^2 \hbar^2}{mq^2} = \frac{n^2 \hbar^2}{mr_B^2} \frac{4\pi\epsilon_0 r_n^2}{q^2} = \frac{n^2 \hbar^2}{mr_n^2} \cdot \frac{r_n}{mv^2} = \frac{n^2 \hbar^2}{m^2 v^2 r_n}$$

$$m^2 v^2 r_n^2 = n^2 \hbar^2$$

$$mvr_n = n\hbar$$

$p_\theta = n\hbar$  is the third Bohr postulate

**Prob. 2.3**

(a) Find generic equation for Lyman, Balmer, and Paschen series.

$$\Delta E = \frac{hc}{\lambda} = \frac{mq^4}{32\pi^2\epsilon_0^2 n_1^2 \hbar^2} - \frac{mq^4}{32\pi^2\epsilon_0^2 n_2^2 \hbar^2}$$

$$\frac{hc}{\lambda} = \frac{mq^4(n_2^2 - n_1^2)}{32\epsilon_0^2 n_1^2 n_2^2 \hbar^2 \pi^2} = \frac{mq^4(n_2^2 - n_1^2)}{8\epsilon_0^2 n_1^2 n_2^2 \hbar^2}$$

$$\lambda = \frac{8\epsilon_0^2 n_1^2 n_2^2 \hbar^2 \cdot hc}{mq^4(n_2^2 - n_1^2)} = \frac{8\epsilon_0^2 \hbar^3 c}{mq^4} \cdot \frac{n_1^2 n_2^2}{n_2^2 - n_1^2}$$

$$\lambda = \frac{8(8.85 \cdot 10^{-12} \frac{F}{m})^2 \cdot (6.63 \cdot 10^{-34} \text{J}\cdot\text{s})^3 \cdot 2.998 \cdot 10^8 \frac{m}{s}}{9.11 \cdot 10^{-31} \text{kg} \cdot (1.60 \cdot 10^{-19} \text{C})^4} \cdot \frac{n_1^2 n_2^2}{n_2^2 - n_1^2}$$

$$\lambda = 9.11 \cdot 10^8 \text{m} \cdot \frac{n_1^2 n_2^2}{n_2^2 - n_1^2} = 9.11 \text{\AA} \cdot \frac{n_1^2 n_2^2}{n_2^2 - n_1^2}$$

$n_1=1$  for Lyman, 2 for Balmer, and 3 for Paschen

(b) Plot wavelength versus  $n$  for Lyman, Balmer, and Paschen series.

LYMAN SERIES				
$n$	$n^2$	$n^2-1$	$\frac{n^2}{n^2-1}$	$911 \cdot \frac{n^2}{n^2-1}$
2	4	3	1.33	1215
3	9	8	1.13	1025
4	16	15	1.07	972
5	25	24	1.04	949

LYMAN LIMIT 911Å

BALMER SERIES				
$n$	$n^2$	$n^2-4$	$\frac{4n^2}{n^2-4}$	$911 \cdot \frac{4n^2}{n^2-4}$
3	9	5	7.20	6569
4	16	12	5.33	4859
5	25	21	4.76	4338
6	36	32	4.50	4100
7	49	45	4.36	3968

BALMER LIMIT 3644Å

PASCHEN SERIES				
$n$	$n^2$	$n^2-9$	$\frac{9n^2}{n^2-9}$	$911 \cdot \frac{9n^2}{n^2-9}$
4	16	7	20.57	18741
5	25	16	14.06	12811
6	36	27	12.00	10932
7	49	40	11.03	10044
8	64	55	10.47	9541
9	81	72	10.13	9224
10	100	91	9.89	9010

PASCHEN LIMIT 8199Å

**Prob. 2.4**

Show equation 2-17 corresponds to equation 2-3. That is show  $c \cdot R = \frac{m \cdot q^4}{2 \cdot K^2 \cdot \hbar^2 \cdot h}$

From 2-17 and solution to 2.3,

$$v_{21} = \frac{c}{\lambda} = \frac{2.998 \cdot 10^8 \frac{m}{s}}{9.11 \cdot 10^8 \text{m} \cdot \frac{n_1^2 n_2^2}{n_2^2 - n_1^2}} = 3.29 \cdot 10^{15} \text{Hz} \cdot \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

From 2-3,

$$v_{21} = c \cdot R \cdot \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right) = 2.998 \cdot 10^8 \frac{m}{s} \cdot 1.097 \cdot 10^7 \frac{1}{m} \cdot \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right) = 3.29 \cdot 10^{15} \text{Hz} \cdot \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

**Prob. 2.5**

(a) Find  $\Delta p_x$  for  $\Delta x = 1 \text{ \AA}$ .

$$\Delta p_x \cdot \Delta x = \frac{h}{4\pi} \rightarrow \Delta p_x = \frac{h}{4\pi \cdot \Delta x} = \frac{6.63 \cdot 10^{-34} \text{ J} \cdot \text{s}}{4\pi \cdot 10^{-10} \text{ m}} = 5.03 \cdot 10^{-25} \frac{\text{kg} \cdot \text{m}}{\text{s}}$$

(b) Find  $\Delta t$  for  $\Delta E = 1 \text{ eV}$ .

$$\Delta E \cdot \Delta t = \frac{h}{4\pi} \rightarrow \Delta t = \frac{h}{4\pi \cdot \Delta E} = \frac{4.14 \cdot 10^{-15} \text{ eV} \cdot \text{s}}{4\pi \cdot 1 \text{ eV}} = 3.30 \cdot 10^{-16} \text{ s}$$

**Prob. 2.6**

Find wavelength of 100eV and 12keV electrons. Comment on electron microscopes compared to visible light microscopes.

$$E = \frac{1}{2}mv^2 \rightarrow v = \sqrt{\frac{2 \cdot E}{m}}$$

$$\lambda = \frac{h}{p} = \frac{h}{mv} = \frac{h}{\sqrt{2 \cdot E \cdot m}} = \frac{6.63 \cdot 10^{-34} \text{ J} \cdot \text{s}}{\sqrt{2 \cdot 9.11 \cdot 10^{-31} \text{ kg}} \cdot E^{\frac{1}{2}}} = E^{-\frac{1}{2}} \cdot 4.91 \cdot 10^{-19} \text{ J}^{\frac{1}{2}} \cdot \text{m}$$

For 100eV,

$$\lambda = E^{-\frac{1}{2}} \cdot 4.91 \cdot 10^{-19} \text{ J}^{\frac{1}{2}} \cdot \text{m} = (100 \text{ eV} \cdot 1.602 \cdot 10^{-19} \frac{\text{J}}{\text{eV}})^{-\frac{1}{2}} \cdot 4.91 \cdot 10^{-19} \text{ J}^{\frac{1}{2}} \cdot \text{m} = 1.23 \cdot 10^{-10} \text{ m} = 1.23 \text{ \AA}$$

For 12keV,

$$\lambda = E^{-\frac{1}{2}} \cdot 4.91 \cdot 10^{-19} \text{ J}^{\frac{1}{2}} \cdot \text{m} = (1.2 \cdot 10^4 \text{ eV} \cdot 1.602 \cdot 10^{-19} \frac{\text{J}}{\text{eV}})^{-\frac{1}{2}} \cdot 4.91 \cdot 10^{-19} \text{ J}^{\frac{1}{2}} \cdot \text{m} = 1.12 \cdot 10^{-11} \text{ m} = 0.112 \text{ \AA}$$

The resolution on a visible microscope is dependent on the wavelength of the light which is around 5000Å; so, the much smaller electron wavelengths provide much better resolution.

**Prob. 2.7**

Show that  $\tau$  is the average lifetime in exponential radioactive decay.

The probability of finding an atom in the stable state at time  $t$  is  $N(t) = N_0 \cdot e^{-\frac{t}{\tau}}$ . This is analogous to the probability of finding a particle at position  $x$  for finding the average.

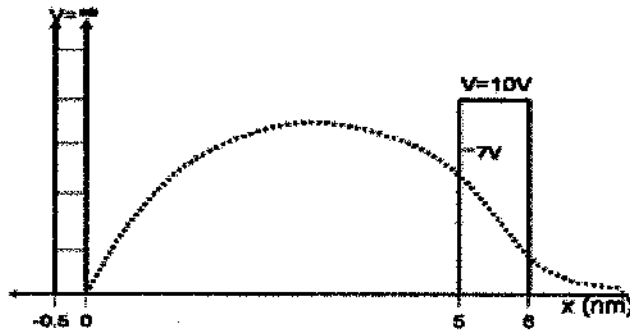
$$\langle t \rangle = \frac{\int_0^{\infty} t e^{-\frac{t}{\tau}} dt}{\int_0^{\infty} e^{-\frac{t}{\tau}} dt} = \frac{\tau^2}{\tau} = \tau$$

This may also be found by mimicking the diffusion length calculation (Equations 4-37 to 4-39).

**Prob. 2.8**

Find the probability of finding an electron at  $x < 0$ . Is the probability of finding an electron at  $x > 0$  zero or non-zero? Is the classical probability of finding an electron at  $x > 6$  zero or non?

The energy barrier at  $x=0$  is infinite; so, there is zero probability of finding an electron at  $x < 0$  ( $|\psi|^2=0$ ). However, it is possible for electrons to tunnel through the barrier at  $5 < x < 6$ ; so, the probability of finding an electron at  $x > 6$  would be quantum mechanically greater than zero ( $|\psi|^2 > 0$ ) and classical mechanically zero.



**Prob. 2.9**

Find  $4 \cdot p_x^2 + 2 \cdot p_z^2 + \frac{7 \cdot E}{m}$  for  $\Psi(x, y, z, t) = A \cdot e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)}$ .

$$\langle p_x^2 \rangle = \frac{\int_{-\infty}^{\infty} A^* \cdot e^{-j(10 \cdot x + 3 \cdot y - 4 \cdot t)} \left( \frac{\hbar}{j} \frac{\partial}{\partial x} \right)^2 A \cdot e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)} dx}{\int_{-\infty}^{\infty} |A|^2 e^{-j(10 \cdot x + 3 \cdot y - 4 \cdot t)} e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)} dx} = 100 \cdot \hbar^2$$

$$\langle p_z^2 \rangle = \frac{\int_{-\infty}^{\infty} A^* \cdot e^{-j(10 \cdot x + 3 \cdot y - 4 \cdot t)} \left( \frac{\hbar}{j} \frac{\partial}{\partial z} \right)^2 A \cdot e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)} dz}{\int_{-\infty}^{\infty} |A|^2 e^{-j(10 \cdot x + 3 \cdot y - 4 \cdot t)} e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)} dz} = 0$$

$$\langle E \rangle = \frac{\int_{-\infty}^{\infty} A^* \cdot e^{-j(10 \cdot x + 3 \cdot y - 4 \cdot t)} \left( -\frac{\hbar}{j} \frac{\partial}{\partial t} \right) A \cdot e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)} dt}{\int_{-\infty}^{\infty} |A|^2 e^{-j(10 \cdot x + 3 \cdot y - 4 \cdot t)} e^{j(10 \cdot x + 3 \cdot y - 4 \cdot t)} dt} = 4 \cdot \hbar$$

$$4 \cdot p_x^2 + 2 \cdot p_z^2 + \frac{7 \cdot E}{m} = 400 \hbar^2 + \frac{28 \hbar}{9.11 \cdot 10^{-31} \text{ kg}}$$

**Prob. 2.10**

Find the uncertainty in position ( $\Delta x$ ) and momentum ( $\Delta p$ ).

$$\Psi(x,t) = \sqrt{\frac{2}{L}} \cdot \sin\left(\frac{\pi x}{L}\right) \cdot e^{-2\pi i E t / \hbar} \quad \text{and} \quad \int_0^L \Psi^* \cdot \Psi dx = 1$$

$$\langle x \rangle = \int_0^L \Psi^* \cdot x \cdot \Psi dx = \frac{2}{L} \int_0^L x \cdot \sin^2\left(\frac{\pi x}{L}\right) dx = 0.5L \quad (\text{from problem note})$$

$$\langle x^2 \rangle = \int_0^L \Psi^* \cdot x^2 \cdot \Psi dx = \frac{2}{L} \int_0^L x^2 \cdot \sin^2\left(\frac{\pi x}{L}\right) dx = 0.28L^2 \quad (\text{from problem note})$$

$$\Delta x = \sqrt{\langle x^2 \rangle - \langle x \rangle^2} = \sqrt{0.28L^2 - (0.5L)^2} = 0.17L$$

$$\Delta p \geq \frac{\hbar}{4\pi \cdot \Delta x} = 0.47 \cdot \frac{\hbar}{L}$$

**Prob. 2.11**

Calculate the first three energy levels for a  $10\text{\AA}$  quantum well with infinite walls.

$$E_n = \frac{n^2 \cdot \pi^2 \cdot \hbar^2}{2 \cdot m \cdot L^2} = \frac{(6.63 \cdot 10^{-34})^2}{8 \cdot 9.11 \cdot 10^{-31} \cdot (10^{-9})^2} \cdot n^2 = 6.03 \cdot 10^{-20} \cdot n^2$$

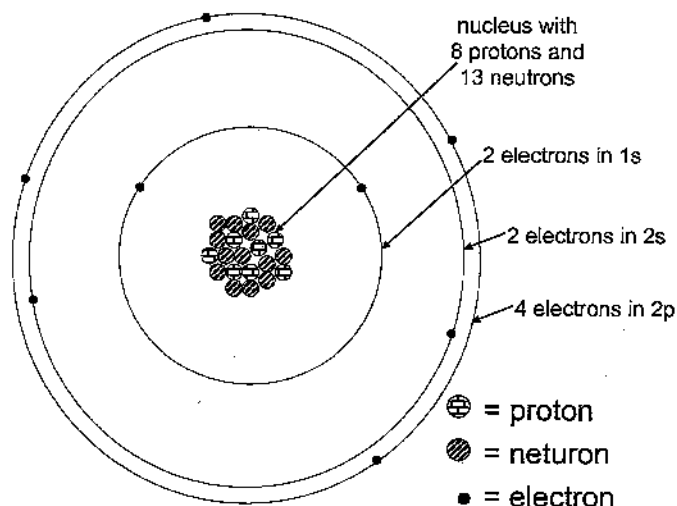
$$E_1 = 6.03 \cdot 10^{-20} \text{J} = 0.377 \text{eV}$$

$$E_2 = 4 \cdot 0.377 \text{eV} = 1.508 \text{eV}$$

$$E_3 = 9 \cdot 0.377 \text{eV} = 3.393 \text{eV}$$

**Prob. 2.12**

Show schematic of atom with  $1s^2 2s^2 2p^4$  and atomic weight 21. Comment on its reactivity.



This atom is chemically reactive because the outer 2p shell is not full. It will tend to try to add two electrons to that outer shell.